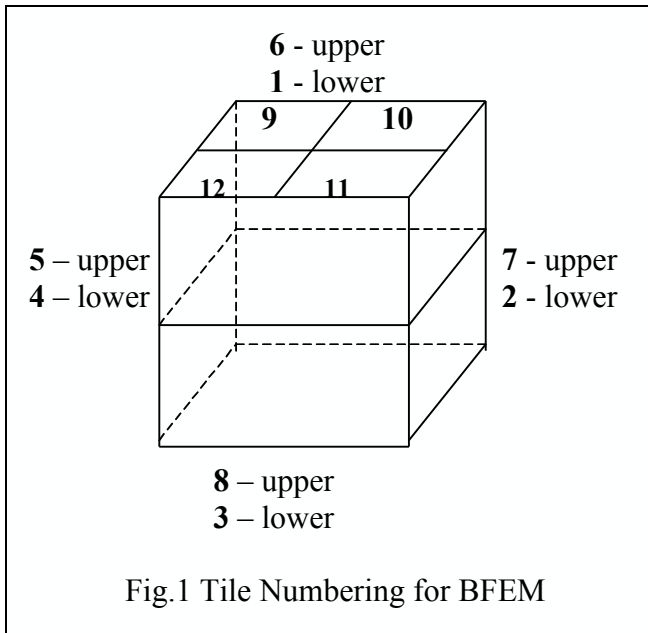


## ACD Gain Calibration Test with cosmic ray muons

**Introduction.** The idea of the test is to run the ACD with cosmic ray muons and obtain pulse-height histograms, corresponding to a MIP, for each tile. These histograms will determine the tile light yield averaged over the tile area with muons distributed uniformly. It was shown (A. Moiseev, Effect of broken fibers on tile efficiency) that the loss of 2-3 (**check!**) fibers will cause the shift of MIP pulse-height peak by 10-15% (**check!**). In testing the flight ACD, the readout will be gated by a VETO signal from ANY tile. The task is to find which tile coincidence combination is best for analyzing each particular tile. It can be done by simulating the ACD with the cosmic ray muon flux. To prove this approach, both simulations and real measurements were performed with BFEM ACD.

### **BFEM Muon test.**



**The test** was performed by pulse height analyzing the signals from each BFEM tile, gated by the signal from one of the tiles. It was repeated twice, first gating with tile 11 (one of the top tiles), and second gating with tile 8 (an upper side tile).

**Simulations** were done by GEANT 3.21/FLUKA with the exact BFEM geometry. The same tiles, 11 and 8, were used for event triggering. The muon flux used for simulations, was taken from A. Stephens (fig.2).

### **Comparison of simulation and test results.**

Table 1 shows the fraction of triggers showing signals in each tile. The pulse-height distributions obtained in the muon test and in the simulations, both triggered by tile 8, are shown in fig. 3 and fig.4, respectively. (Similar results triggered by tile 11 are not shown here.) Comparison between simulation and test results show that they are consistent, and that the simulations can be used to develop the muon test technique for the flight ACD configuration.

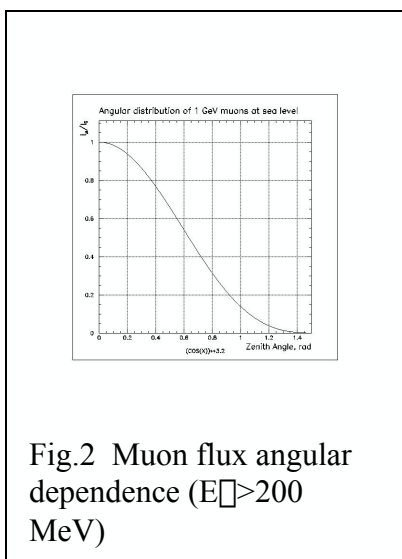


Table 1. Fraction of triggers accompanied by a signal in each BFEM tile

Tile number	Test – gated by 11	Simulation – gated by 11	Test – gated by 8	Simulation – gated by 8
1	0.04	0.043	0.049	0.056
2	0.055	0.068	0.049	0.061
3	0.067	0.079		
4	0.049	0.042	0.052	0.062
5	0.065 ?	0.018	0.068	0.083
6	0.022	0.018	0.028	0.023
7	0.09	0.17	0.066	0.084
8	0.21	0.17		
9			0.018	0.019
10			0.017	0.018
11			0.11	0.18
12			0.11	0.17

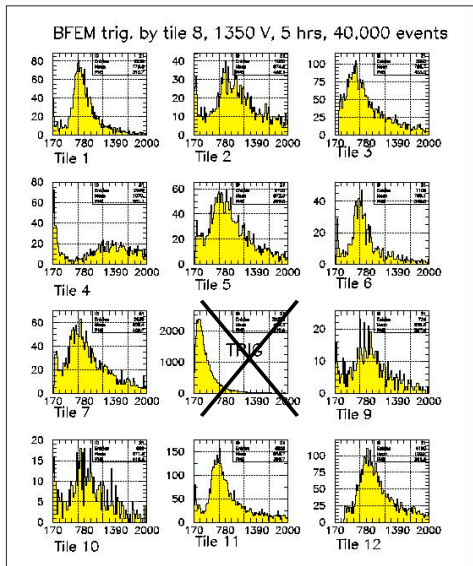


Fig.3 Muon test: Histograms from BFEM tiles triggered by tile 8

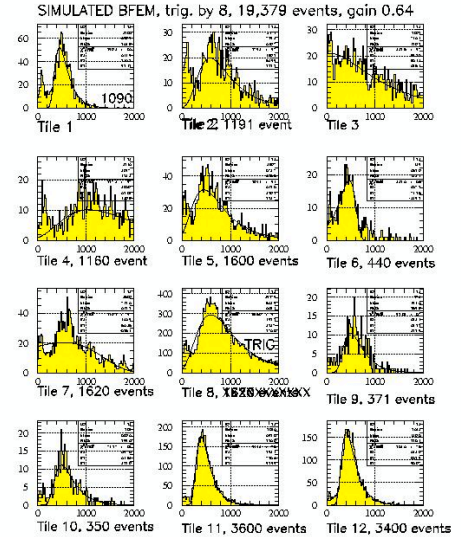
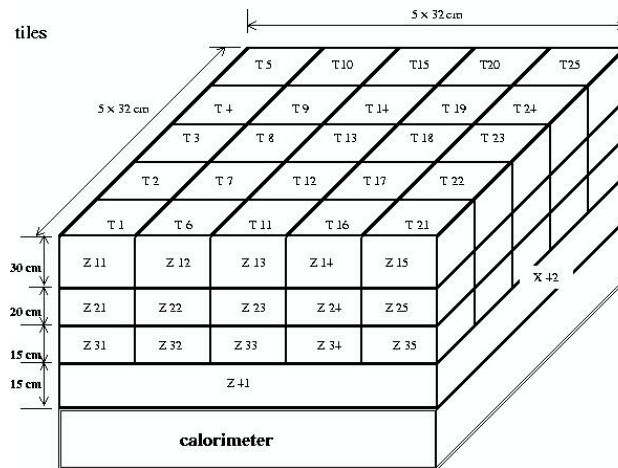


Fig.4 Simulations: Histograms from BFEM tiles triggered by tile 8

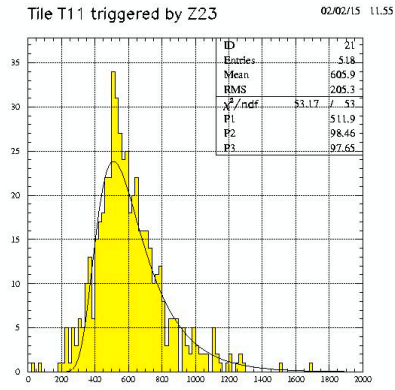
**Test of the Flight ACD**. The task is to find for each particular tile what tiles can be used to trigger it. The trajectories should be as normal to the tested tile surface as possible, with reasonable statistics to be collected from the cosmic muons. The tile numbering used in the simulations is shown in Fig. 5. The simulation run corresponded to approximately 40 minutes of ACD running time, with 10-11 thousand triggers



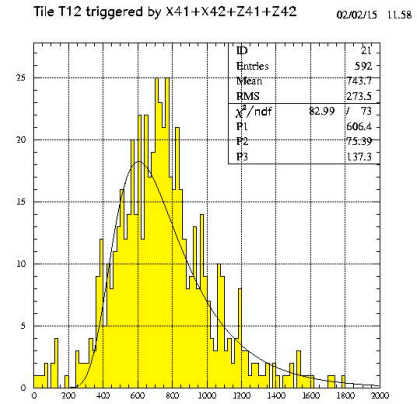
**Fig. 5**

collected for each of top tiles. For each tile, the triggering tiles were carefully selected, and corresponding histograms are shown in figures below. For reliable fitting and MIP peak position determination, approximately 1,500 events are desirable in the histogram. Looking at the histograms, we see that the most difficult tiles to calibrate will be the upper side tiles (fig. 7), which will require 6-7 hours to get  $\sim 1,500$  events. Limited calibration can be done within  $\sim 4$  hours.

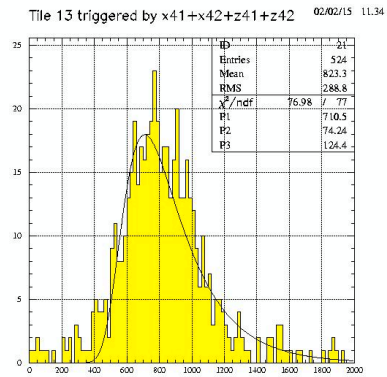
**Fig.6 - Top tiles**



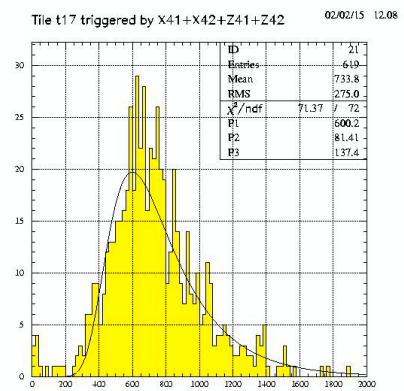
Tile T11 – middle of top edge



Tile T12 – next to the top central tile

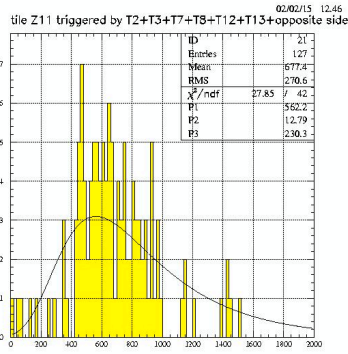


Tile T13 – top center tile

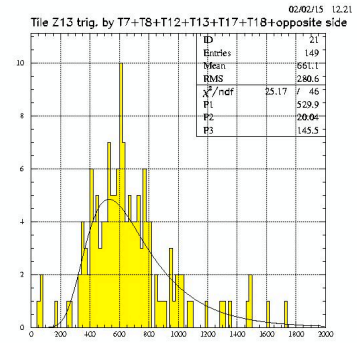


Tile T17 – diagonal from the top central tile

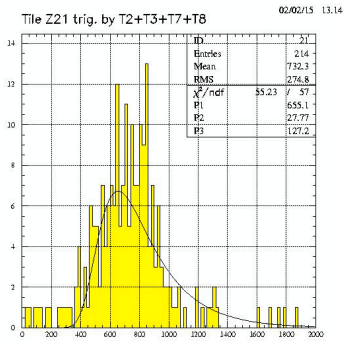
**Fig. 7 - Side tiles**



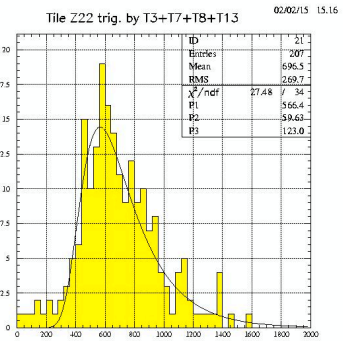
Tile Z11 – end tile in the side top row



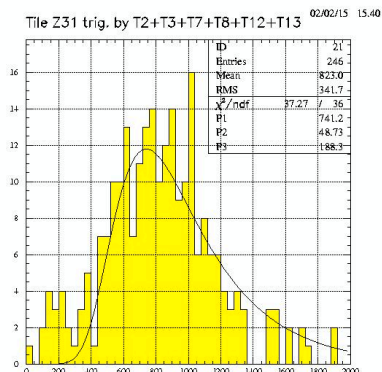
Tile Z13 – middle tile in the side top row



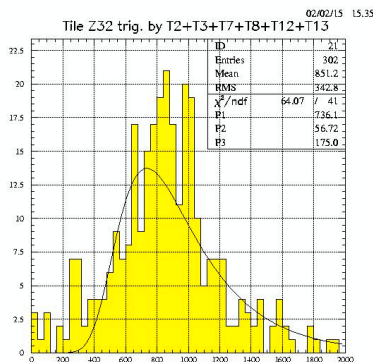
Tile Z21 – end tile in side second row



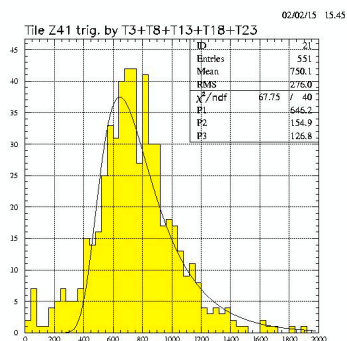
Tile Z22 – second tile from end in side second row



Tile 31 – end tile in side third row



Tile Z32 – second tile from end in side third row



Tile Z41 - Long tile, side bottom row

### Some remarks how to use these histograms:

1. Look at the quality of the histogram. For those that have too few of events, project how would it look at higher statistics.
2. These histograms would be collected in ~40 min of running time, so one can estimate how much time will be needed to get given number of events.

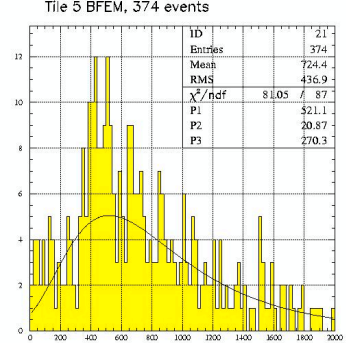
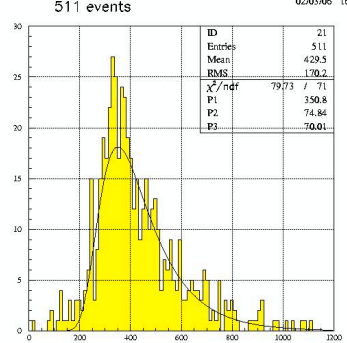
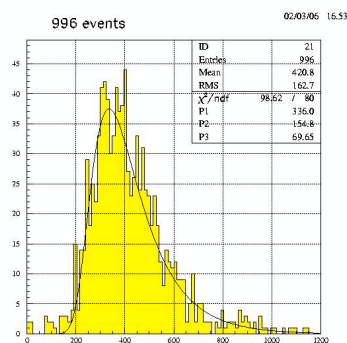
Note which tiles were used for selecting events in each histogram. Any suggestions to improve these choices would be welcome.

**How many events do we need in the histogram for given peak position uncertainty?**

I believe that the common mathematical approach to this estimate is complicated by the high variability of the particle paths in the tile and desirable lowering the number of events needed. The simulations seem to be the appropriate way to do this analysis. I did the following – using sea level muon flux in the simulations, the simulated pulse height distribution was fitted by Landau distribution to find the peak position. This was repeated for 10 sets of approximately 2,500, 1,000, and 500 events in the histogram and the mean value and standard deviation ( $\sigma$ ) was determined for each of these sets of 10 runs. This was done for the tile on the top of ACD (with the most of muons hit the tile on or around normal incidence). The results are given in Table 2. The examples of the pulse height distribution for the top tile, 996 events and 511 events in the histogram are given in fig. 8 and fig.9 (column 3 and 4 in the table) respectively. The histogram for the side (“bad”) tile and 374 events is shown in fig. 10 (5-th column in the table). It is seen that the precision of the peak position fitting is surprisingly high, even for such a small statistics as  $\sim 500$  events are and for the “bad” tile (the side one) there are a large variation of the incident muon angles, and consequently the muon paths in the tile. Let me remind that we are looking for the change in the light yield on a level of 5% and more.

**Table 2** Simulations of the peak position determination precision

Fitted MIP peak position	Top tile, $\sim 2,500$ events	Top tile, $\sim 1,000$ events	Top tile, $\sim 500$ events	Side tile, $\sim 400$ events, gain=1	Side tile, $\sim 400$ events, gain=0.95
	330.8	329.9	332.2	487.3	467.4
	333.0	326.8	338.2	495.2	467.2
	336.7	336.6	324.5	510.3	462.8
	340.9	338.4	331.5	481.3	489.9
	330.7	335.0	340.9	482.1	478.9
	334.9	330.9	348.4	491.9	539.5 $\sigma$
	331.0	336.0	326.8	509.5	447.1
	336.8	339.1	343.0	510.9	482.1
	331.2	339.5	336.5	511.7	482.5
	336.2	334.8	350.8	521.1	476.4
Mean $\pm \sigma$	334.2 $\pm$ 3.4 (1%)	334.7 $\pm$ 4.3 (1.3%)	337.3 $\pm$ 8.7 (2.6%)	500.1 $\pm$ 13.8 (2.8%)	478.7 $\pm$ 24 (5%)



**Fig. 8** Top tile, 996 events

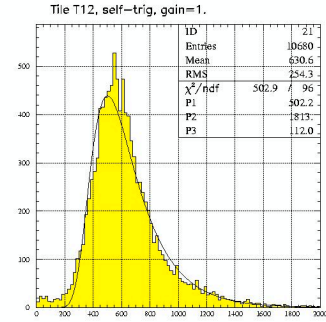
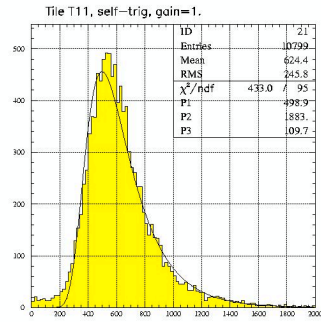
**Fig.9** Top tile, 511 events

**Fig. 10** side tile, 374 events

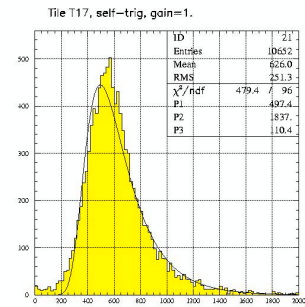
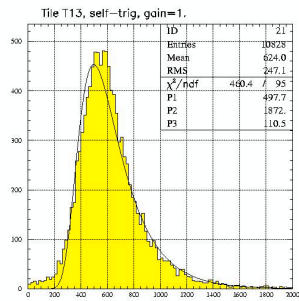
**The back-up option of testing flight ACD.** Another way to do the gain calibration test is to look at all tile histograms in muon run self-triggering mode, meaning that all signals appeared in given tile, will be used for the analysis. The advantage of this approach is that for  $\sim 1$  hour of instrument running there will be from 3,000 to 15,000 events in the histogram (depending on the tile) which provides a very reliable and precise peak position determination. The disadvantage of this approach is that the histograms for some tiles, especially for the side tiles, will be very dependable on the muon flux angular distribution. This is because no external triggering will be used, so the angular range of particles causing the triggering will be  $2\pi$  for every tile, which in the convolution with the incident flux angular dependence could cause the uncertainty if the latter one varies. I believe that the muon flux angular distribution is constant for the given place of measurements, so this particular approach can be successfully used for the repeated functional tests performed in the same place. After moving to the other place, the re-calibration should be done using the approach described earlier. Similarly to fig. 6 and 7 with histograms selected by appropriate triggering, the histograms for the same tiles in self-triggering mode, are presented in fig. 11 and 12. We see how high statistics is there.



**Fig. 11** – Top tiles in self-triggering mode



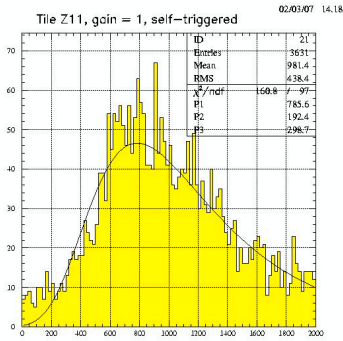
Tile T11 – middle tile in the edge row      Tile T12 – next to the central tile



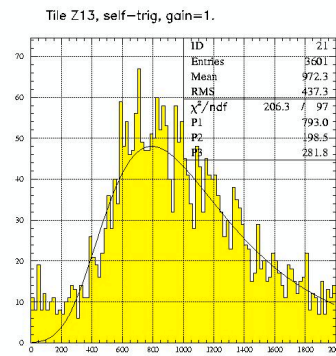
Tile T13 – top center tile

Tile T17 – diagonal from the central tile

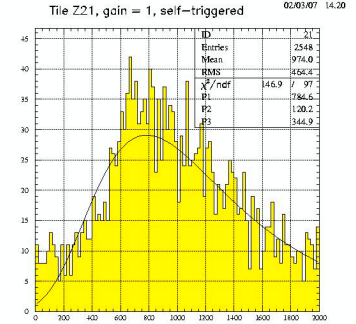
**Fig. 12 – Side tiles**



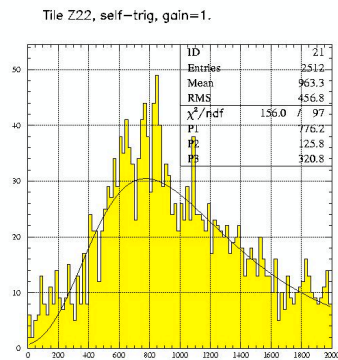
Tile Z11 – end tile in the top row



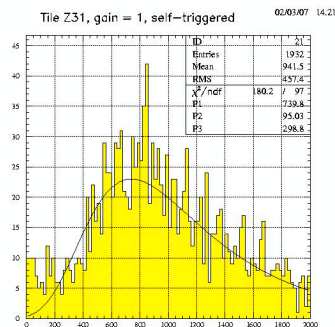
Tile Z13 – middle tile in the top row



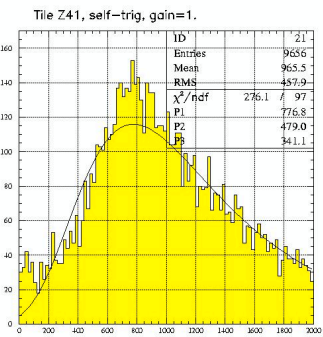
Tile Z21 – end tile in the second row



Tile Z22 – second tile from end in side second row



Tile Z31 – end tile in 3-rd row



Tile Z41 – Long tile bottom row

**How sensitive is the self-triggering mode to the gain change?** The gain change by 5% was simulated. Similar thing was experimentally tested and proven with muons on BFEM. The simulation run identical to that given above, but with the gain of 0.95 was performed. The MIP peak positions for these runs are given in the Table 3 along with the statistics in each corresponding histogram.

Tile	Statistics for 40 min	Peak position for gain = 1	Peak position for gain = 0.95	Ratio
T11	10,650	498.9	472.7	0.947
T12	10,680	502.2	476.4	0.949
T13	10,828	497.7	471.6	0.947
T17	10,651	497.4	471.6	0.948
Z11	3,631	785.6	748.9	0.953
Z13	3,601	793.0	756.7	0.954
Z21	2,547	784.6	748.8	0.954
Z22	2,512	776.2	729.7	0.940
Z31	1,930	739.8	711.2	0.961
Z32	1,944	765.5	735.8	0.961
Z41	9,658	776.8	738.0	0.950

Obtained results demonstrate that the approach is quite sensitive on the required level of sensitivity (5% of the gain relative change)

**Conclusion.** I believe that the gain calibration test should be done as follows:

1. During ACD I&T – a) gain calibration with muon hodoscope for each tile. This will be the most precise measurement (almost free from the uncertainty introduced by the different muon arrival direction). These results will serve as reference in a case of unclear future test results  
b) muon run for 8-10 hours with triggering from any tile. The results will be treated by two ways – selecting triggers, and in the self-triggering mode. In both the statistics reduction will be used to understand the stability of the results
2. In all other ACD test – the muon run will be used for the allowed time, and depending on the time the approach will be used. In the case of visible performance change more careful test should be performed, possibly requiring more time. In some extreme cases, when there will be a possibility of tile replacement, the muon hodoscope test should be performed before making a decision to replace the tile.